Ellipsometer

SOP

The J. A. Woollam M88 is a spectroscopic ellipsometer used to measure film thickness and optical constants of transparent/semi-transparent thin films. It uses a Xenon arc lamp source and measures spectroscopic data over 88 discrete wavelengths from 280-763nm. The tool also offers variable angle of incidence and a spot size adjustment aperture (spot size from ~2cm to 5mm).

Safety

• **Buddy System** – the Buddy System is not needed for after-hours usage of this equipment

1.0 Pre-Operation

1.1 Log in to TUMI system

Note: For the best accuracy, let the lamp warm up for 30 minutes. Start the lamp prior to your TUMI login so you won't be charged for warmup time.

2.0 Restrictions

2.1 None

Ellipsometer Components





Operation

3.0 Startup

- 3.1 Turn on power to the light source LPS-300 and ignite lamp. Note: For the best accuracy, let the lamp warm up for 30 minutes. Start the lamp prior to your TUMI login so you won't be charged for warmup time.
- 3.2 Turn on EC110 control module.
- 3.3 Click the WVASE32 icon to start the software.
- 3.4 There are six different windows (Model, Hardware, Generated Data, Experimental Data, Fit, and Graph) that can be active and you will need to utilize most of them. The menu bar options will be different depending on which window is active.
- 3.5 Frist, select the **Hardware** window and then click on "Initialize" in the menu bar.
- 3.6 Enter a user name when prompted.

4.0 Calibration

- 4.1 Place the silicon calibration wafer on the stage.
- 4.2 Click on "Acquire data \rightarrow Align sample" from the menu bar.
 - 4.2.1. Adjust the stage tilt knobs until the red crosshairs are centered as shown below
 - 4.2.2. Press the escape (Esc) button to exit



- 4.3 Click on "Acquire data \rightarrow Calibrate system" from the menu bar.
 - 4.3.1. Select "Normal" and ensure that the angle of incidence is 75° and click "OK".

Calibration Routine	E Constantino de la C
Calibration Mode	1
C llormat +	_
C Coarse (requires a but	k-like' sample)
C Detector Electronics (of	ff Sample)
C Detector System (Straig	ght-through configuration)
C Detector System (with	Calibration Sample)
Revs per Measurement:	10
# of <u>P</u> oints:	20
Degree Span (of polarizer):	50
Angle of Incidence:	75 <u>Ok</u>
Include window effects	Cancel

- 4.3.2. It will ask to verify that the angle micrometer is at 35mm (75°). Look at the angle adjustment knob (shown in the component overview) to verify
- 4.3.3. The resulting graph shape should be similar to the one below. Contact NRF staff if it fails the calibration.



5.0 Sample Scanning

- 5.1 When complete, remove the silicon standard and place your sample to be measured on the stage.
- 5.2 Again, click on "Acquire data \rightarrow Align sample" from the menu bar.
 - 5.2.1. Center the red crosshair by using the knobs on the stage as before. Always align a sample when moving or replacing it.

Notes:

1) The <u>entire</u> light beam must be within the sample. For small samples, close the iris until the beam is contained within the extents of the sample.



- 2) For very thick samples, you may need to adjust the height by using the height control knob so that the beam falls on the surface.
- 5.3 Click on "Acquire data→Spectroscopic scan" from the menu bar and a new window will appear.

Wavelength:	All 88 Wavelength(s) are Selected for Data Collec	tion Wy Select
Angle:	E to 75	by 5
- Ellipsometri	ic Data Acquisition Setting	
Ren	rs/meas: 5	
Polarizer	setting: 30	<u>O</u> k
FH	igh Accuracy Node	Cancel
	5	1
[Iranin	ion Maarman	

5.3.1. Most measurements can be taken at a fixed angle of 75°, and all of the default settings shown above.

Note: For difficult samples and when fitting optical constants, multiple angles may be needed. For multi-angle measurements, set starting angle in first box, final angle in second box and the degree increments in the third box. You will need to adjust the angle micrometer when prompted.

Note: Support the polarizer or detector arm when changing angles. Failure to do so can result in damage to the micrometer.

- 5.3.2. Ensure all 88 wavelengths are selected and click "OK"
- 5.4 The scan is extremely fast and the graph will now be populated with the experimental data (Psi and Delta)
- 5.5 It may be beneficial to turn off the double axes because the data can get very difficult to analyze as shown in the graph on the left. Single axis (right graph) will make model fitting much easier.



5.6 To turn off double axes, select the **Graph** window. Click on "Style→2D" and then remove the check mark for double Y-axis

2D Graph Style
☑ Double Y-axis
Graph as function of Angle
🗆 Trajectory plot (dynamic data)
Exchange Translation Coordinates
Reverse Axis: □X □Y1 □Y2
🗆 Plot Derivatives, 🛛 Order: 🔟
of Points: 10
Polynomial degree: 3
Ok Cancel

6.0 Modelling And Fitting Thickness (single layer, tabulated model)

- 6.1 Select the Model window
- 6.2 You will need to build a model that represents your sample by the layers that are present. Think of it as looking at your sample in cross-section. Build your model by selecting "Add layer" in the menu bar. The window that opens will contain some commonly used materials. More materials are located in the subfolders and organized by material type.

Open			? 🛛
File name: <mark>*.mat</mark>	Folders: c:\wvase32\mat		OK
7059_g.mat Acetone_g.mat Ag_g.mat a-GaAs_Adachi.mat a-GaAs_cl.mat a-GaP_Adachi.mat a-Ge_Adachi.mat a-Ge_Adachi.mat	Archive Dielectr IIR_Mats Metals Semicond Special	 	
List files of type: Default Materials	Drives:	_	Network

Note: Most material files are tabulated lists of index of refraction/extinction

coefficient (n&k) – vs – wavelength. However, n&k can vary with film deposition conditions (or post processing) and your film may not match a specific material file, or your material may not be in the database. In this case, you will need to use a function based model. This will be explained in following section.

6.2.1. The following assumes you are using tabulated models. Start with the substrate first and select it from the materials in the list (Silicon shown as an example).

Layer	×
Layer Name: si_jaw	
Comment: Si substrate, Herzinge	r et.al., JAP v83p3323y19
Spectral range of optical constants:	187.88 - 6200 nm
Thickness: 🚺	mm Fit
Optical Constants >> Opt Const Fit	Ok Cancel Replace Laver
Reset n&k	Delete Layer

- 6.2.1.1. If the substrate is non-transparent, then the thickness does not matter (although some materials may become transparent near the IR or UV regions...consult with staff if unsure). Definitely, if it is transparent in the visible then you will need to know the thickness of your substrate
- 6.2.2. Select "Add layer" for the next material (SiO2 shown for example). Enter your initial guess for the thickness now and check the "Fit" box

ayer	X
Layer Name: sio2_jaw Comment: Thermal SiO2, Herzinger et.al., JAP v83p3323	ly1
Spectral range of optical constants: 187.88 - 1771.4 nm Thickness: 100 Å 당 Fit	
Optical Constants >> Ok Opt Const Fit Cancel In Cancel Ik Replace Layer Reset n&k Delete Layer	r

Do not check the n&k boxes!

- 6.3 Select the Generated Data window.
- 6.4 Click on "Generate data" in the menu bar. This will populate the graph

with the modeled data and will be shown in red. In the example below, the model (red) does not fit the experimental data (green) very well. This indicates that the thickness guess is off.



- 6.5 Select the **Model** window and click on the film layer. Adjust the thickness up/down.
- 6.6 Select the **Generated Data** window again, and click on "Generate data". Observe if the two data sets begin to converge or not. Repeat until the two sets are closely fit
- 6.7 Select the **Fit** window.
- 6.8 Click on "Normal Fit" in the menu bar. This will automatically adjust any parameters that you have selected to fit. In this case, it only adjusted thickness. The data should now overlap if the material file constants are close to your film.



6.8.1. Check that the measurement makes sense and that the MSE is small.

7.0 Modelling And Fitting Thickness And Optical Constants (single, transparent layer, Cauchy model)

7.1 When fitting optical constants, it is highly recommended to take

measurements at multiple angles.

7.2 The <u>Cauchy model</u> is the most commonly used approach to determine the thickness and index of <u>transparent materials</u>. It is an inverse power series that contains only even terms and creates the dispersion model for the index of refraction only.

$$n(\lambda) = A + \frac{B}{\lambda^2} + \frac{C}{\lambda^4},$$

- 7.3 Select the **Model** window and load a substrate material, the same as above.
- 7.4 Again, select "Add layer" and load the Cauchy.mat file. Specify a value for the "A" term that roughly matches the expected index value and acts as an offset value. The "B" and "C" terms add curvature. Also specify the expected thickness.

	Cauchy Layer, with Urbach Absorption			ach Absorption
	Layer Name	cauchy		
	Comment: [Cauchy materia	ıl	
	<u>T</u> hickness: [<mark>0.</mark> nr	Ē	t
Cauchy model	<u>A</u> n:	1.45	🗆 Fit	<u>0</u> k
parameters	<u>B</u> n:	1.e-002	🗆 Fit	Delete Layer
	<u>C</u> n:	0.	🗆 Fit	Ontical Constants >>
	k A <u>m</u> plitude:	0.	🗆 Fit	
	Exponent:	1.5	🗆 Fit	
	Band Edge:	4000.	🗆 Fit	
	12		1	

- 7.5 Select the **Generated Data** window, and click on "Generate data". Observe the graph and adjust thickness→Generate data until the model and experimental data are close.
- 7.6 In the **Model** window, click on the Cauchy layer again. This time, check the fit boxes for thickness, A, and B.
- 7.7 Select **Fit** \rightarrow Normal fit and the two curves should converge.
- 7.8 Observe MSE, then open the Cauchy layer again and check the fit box for the "C" term and type a small positive number in the box.
- 7.9 Select **Fit** \rightarrow Normal fit and see if the MSE changes significantly.
- 7.10 To observe optical constants, open the Cauchy layer, then click on the optical constants button and you will see something similar to the

graph below. Verify normal dispersion! See pages 295-298 of the manual for interpreting if the index is valid.



8.0 Modelling And Fitting Thickness And Optical Constants (single, partially absorbing layer, Cauchy model)

- 8.1 When fitting optical constants, it is highly recommended to take measurements at multiple angles.
- 8.2 For partially absorbing materials, you will need to divide the analysis into two parts.
 - 8.2.1. Start with the transparent region where the Cauchy model is valid and determine thickness and index.
 - 8.2.2. Then fix the thickness and fit *n* & *k* over then entire range using point by point fitting.
- 8.3 Select the **Model** window and load a substrate material, the same as above.
- 8.4 Again, select "Add layer" and load the Cauchy.mat file. Specify a value for the "A" term that roughly matches the expected index value. Also specify the expected thickness.
- 8.5 Select the **Generated Data** window, and click on "Generate data". Observe the graph and adjust thickness→Generate data until the model and experimental data are close.
- 8.6 In the **Model** window, click on the Cauchy layer again. This time, check the fit boxes for thickness, A, and B.
- 8.7 Select **Fit**→Normal fit and the two curves should converge in the transparent region. Shown below is a data of a photoresist at three

different angles. The data from the transparent region matches the model, but it does not match in the absorbing region.



8.8 To focus on the transparent region only, select the **Experimental Data** window, the click on "Range Select". Input the wavelength range the corresponds to the transparent region of your film

Select Experimental Data Range	X
Angle: [50] to 90.	
Wavelength: 124.00 to 50000000 hm	
Selected Data Types:	
E: Ellipsometric dpolE: Ellipsometric Depolarization	
China data	
All Data every 0. point(s)	
Add to Currently selected data Ok	
Remove from Currently selected data Cancel	
Global (select range for data associated with all models)	
Range select generated data	

8.9 Select the **Generated Data** window, and click on "Generate data". The graph will now update with the range that you specified.



- 8.10 Select **Fit**→Normal fit (with thickness, A, B, maybe C to be fit) and it should be a much better fit since the model is now only using the data in the transparent region, and is not being influenced by absorption.
- 8.11 We now need to fit n & k over the entire range. Select **Experimental Data** \rightarrow Range Select. Now click the "All Data" button.
- 8.12 Select the **Generated Data** window, and click on "Generate data". The graph will now update with all of the data again.
- 8.13 In the **Model** window, click on the Cauchy layer again. Now deselect all of the thickness and Cauchy parameters, and select the n & k boxes.

Cauchy Layer	, with Urbach Al	bsorption	
Layer Name:	cauchy		
Comment: [Cauchy material		
Thickness: [1619.91 Å	🗆 Fit	Ok
An:	1.5606	🗆 Fit	Delete Layer
Bn:	0.01169	🗖 Fit	Replace Layer
Cn:	0	🗆 Fit	Antical Constants >>
k Amplitude:	0	🖂 Fit 👘	
Exponent:	1.5	🖂 Fit	
Band Edge:	400	nm	₩ k

8.14 Select **Fit**→Pt. by Pt. fit and you will get another window. The only adjustment you need to make is the starting position. You need to start the fitting process where the model overlays the experimental data. In the example, that was at the longest wavelength. Click OK and it will ask again if you want to change the starting conditions, click No.

Wvl by Wvl Fit: Starting Parameter Values 🛛 🔀
n.1698.6.1 <mark>1.5646</mark>
k.1698.6.1 0 Ok
Starting Position
⊂ Shortest WvI © Longest WvI
Pts. to Ave. for next quess: 1.
Show Iterations

8.15 It will now adjust n & k at every wavelength and the fit should be very good, as shown in the example below.



Note: A good fit does not mean it is a valid result. This method can result in physically impossible values for n & k if your model is not valid, or is missing something. Verification is difficult and requires the use of a consistency check. See pages 175-177 for details on using the Kramers-Kronig layer (KK.mat).

8.16 To view the optical constants, open the Cauchy model layer and click on optical constants. To save the constants, click the copy to clipboard button. Open Notepad and paste, then save as a text file. This can be opened in Excel or other graphing software.



9.0 Saving And Removing Data

- 9.1 The experimental data can be saved as just the raw data, or can include the model as well.
- 9.2 To save just the experimental data, select **Experimental Data**→File→Save Exp. File
 - 9.2.1. Make a folder in the User Data folder and type a file name to save it.
- 9.3 To save the entire session (all windows and model info), click on "Global" located in the upper right corner, and select Save Current Environment



- 9.3.1. Again, make a folder in the User Data folder and type a file name to save it.
- 9.4 To remove data from the computer, you need to save it to the floppy disk that is provided
 - 9.4.1. Remove the disk from the main computer and place it in the floppy drive of the offline computer.
 - 9.4.2. The data can now be saved to a USB drive. The USB port is located at the bottom of the offline computer.
- 9.5 If you would only like to save the graph as a picture, simply right click on the graph and select copy to clipboard. You can now open Paint and paste it and save just the picture.

10.0 Shut Down

- 10.1 Exit the WVASE software
- 10.2 Turn off the controller module
- 10.3 Turn off the lamp power supply
- 10.4 Log off the TUMI